The weak interaction Part II



Marie-Hélène Schune Achille Stocchi LAL-Orsay IN2P3/CNRS

- The $K^0-\overline{K}^0$ system
- The CKM mechanism
- Measurements of the unitarity triangle parameters : some examples
- Neutrinos



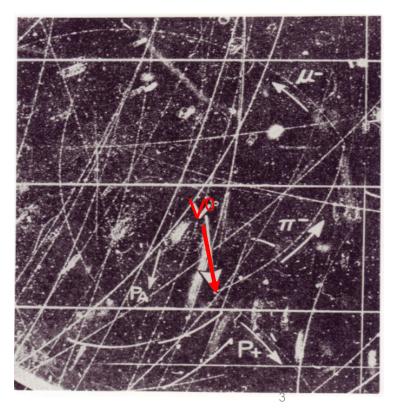
Remember the strange particles ?

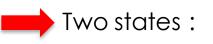
Observation of Long-Lived Neutral V Particles*

K. LANDE, E. T. BOOTH, J. IMPEDUGLIA, AND L. M. LEDERMAN, Columbia University, New York, New York

AND

W. CHINOWSKY, Brookhaven National Laboratory, Upton, New York (Received July 30, 1956)







Brookhaven, 1956

 $M(\pi) \sim 140 \text{ MeV}$

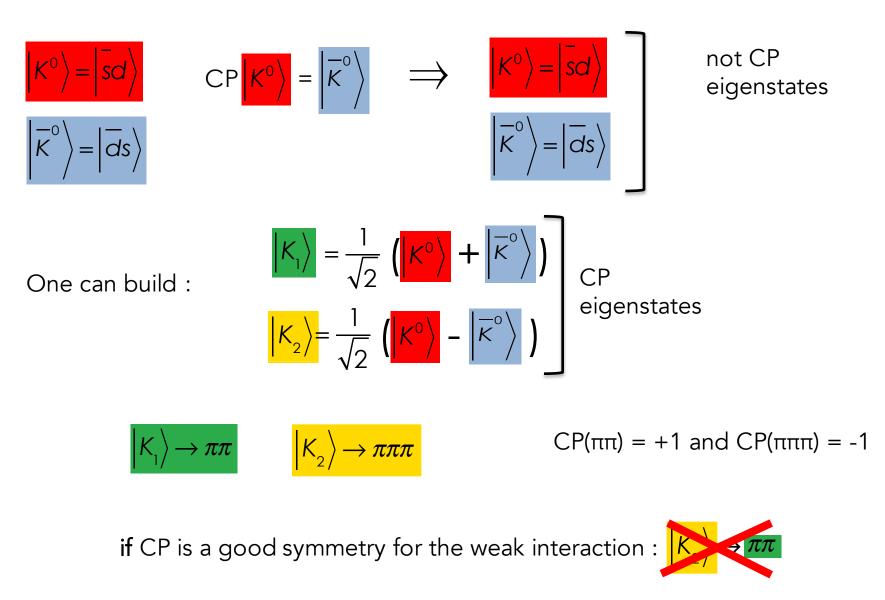
M(K) ~ 500 MeV

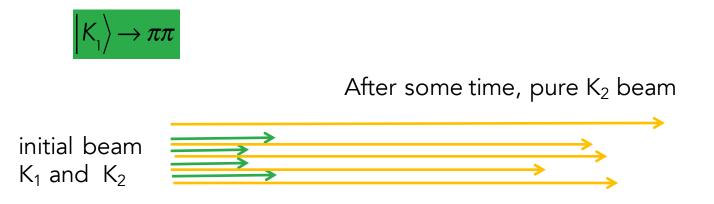
Same mass (~ 500 MeV)

Very different lifetimes

 K_2 lifetime ~ 10000 K_1 lifetime due to phase space

CP violation in the K⁰ system

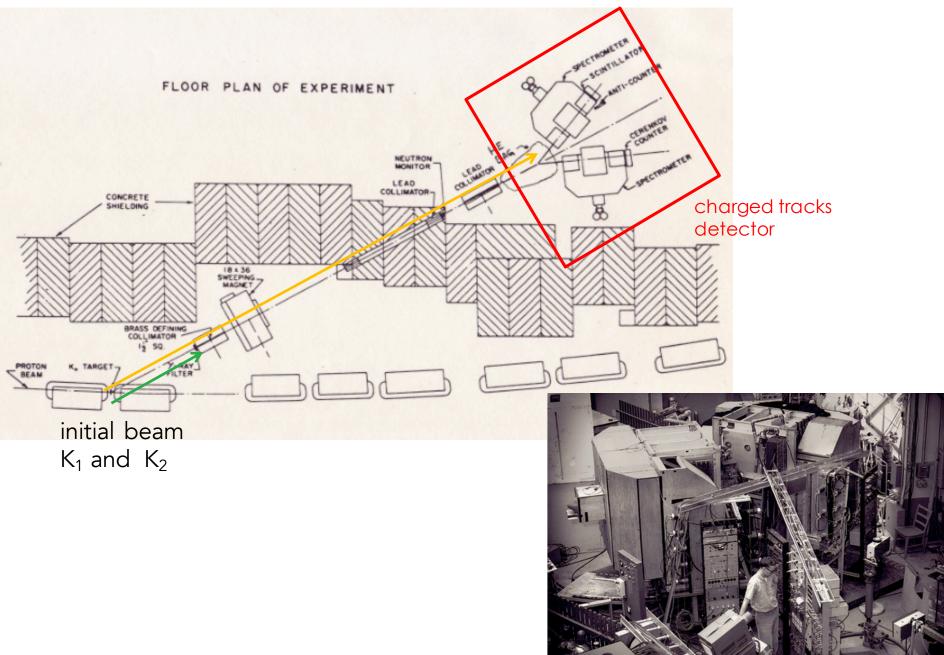




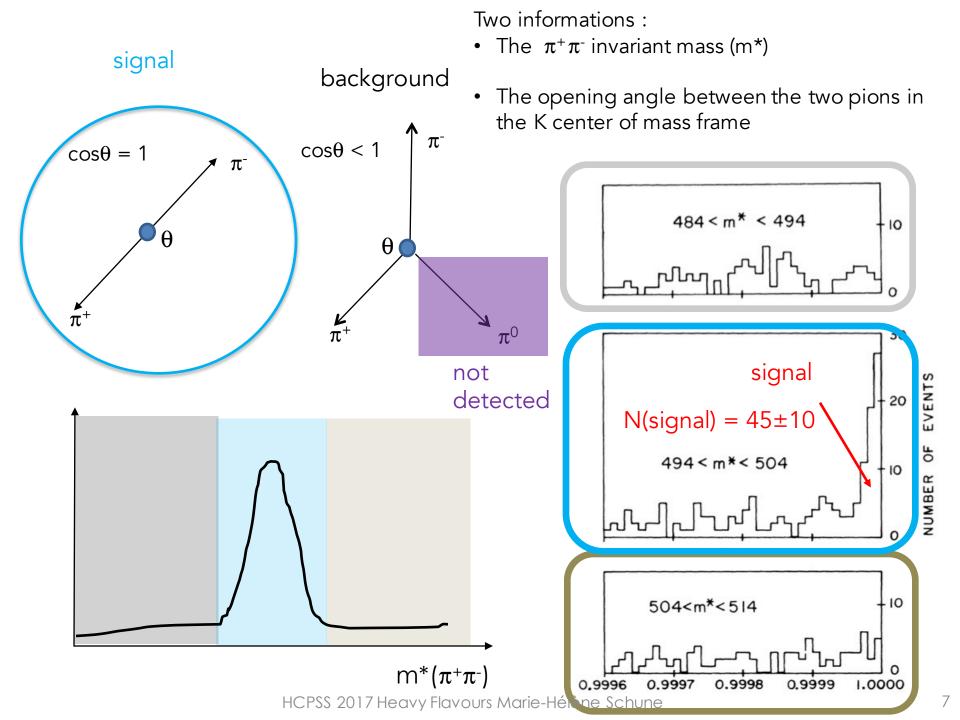
Search for the signal of the decay $|\kappa_2\rangle \rightarrow \pi\pi$ far (20 meters) from the production point of the K_1 and K_2

Ś

Cronin& Fitch experiment 1964



Weak Interaction, An-Najah National University, Nak



1964

We would conclude therefore that K_2^0 decays to two pions with a branching ratio $R = (K_2 - \pi^+ + \pi^-)/(K_2^0 - \text{all charged modes}) = (2.0 \pm 0.4) \times 10^{-3}$ where

the error is the standard deviation. As empha-

sized above, any alternate explanation of the ef-

fect requires highly nonphysical behavior of the

three-body decays of the K_2^{0} . The presence of a two-pion decay mode implies that the K_2^{0} meson

is not a pure eigenstate of CP. Expressed as

EVIDENCE FOR THE 2π DECAY OF THE K^o MESON^{*†}

J. H. Christenson, J. W. Cronin,[‡] V. L. Fitch,[‡] and R. Turlay[§] Princeton University, Princeton, New Jersey (Received 10 July 1964) 27 JULY 1964

The Nobel Prize in Physics 1980





James Watson Cronin Prize share: 1/2

Val Logsdon Fitch Prize share: 1/2

The Nobel Prize in Physics 1980 was awarded jointly to James Watson Cronin and Val Logsdon Fitch "for the discovery of violations of fundamental symmetry principles in the decay of neutral K-mesons"

R. Turlay was a PhD student J Christenson was a graduate studen

« The discovery emphasizes, once again, that even almost self evident principles in science cannot be regarded fully valid until they have been critically examined in precise experiments. »

Today :

$$\frac{A|K_2}{A|K_2} \rightarrow \pi\pi \qquad = \frac{1}{2} (2.271 \pm 0.017) 10^{-3} \qquad 0.7 \ \% \ \text{precision !}$$

Experimental observation of CP violation in K decays

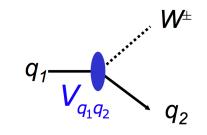
+ Cabibbo angle



V_{CKM} Cabibbo-Kobayashi-Maskawa matrix

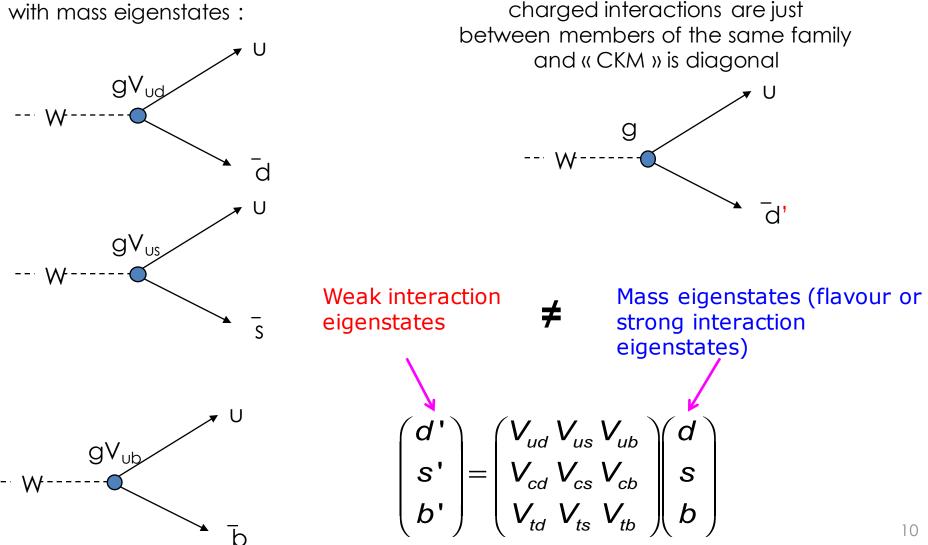
V_{CKM} Cabibbo-Kobayashi-Maskawa matrix

Two different way of seeing the charged interactions among quarks



In the basis where :

In the basis dealing with mass eigenstates :



Progress of Theoretical Physics, Vol. 49, No. 2, February 1973

CP-Violation in the Renormalizable Theory of Weak Interaction

Makoto KOBAYASHI and Toshihide MASKAWA

Department of Physics, Kyoto University, Kyoto

1973 Before the discovery of the 4th quark

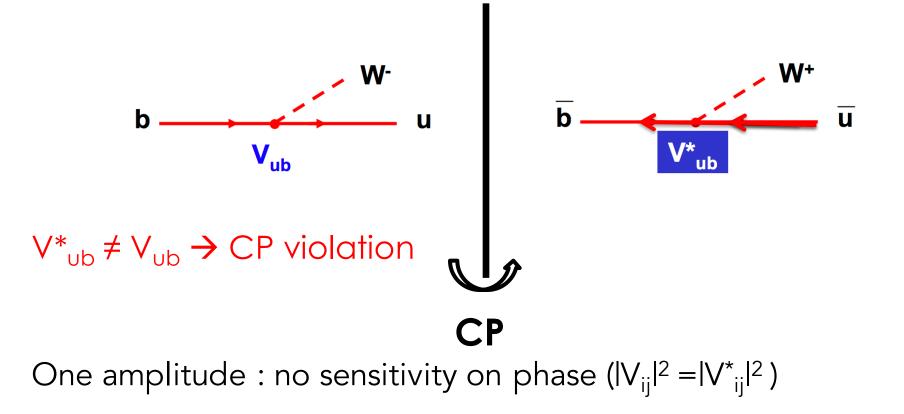
Prediction of the 3rd family

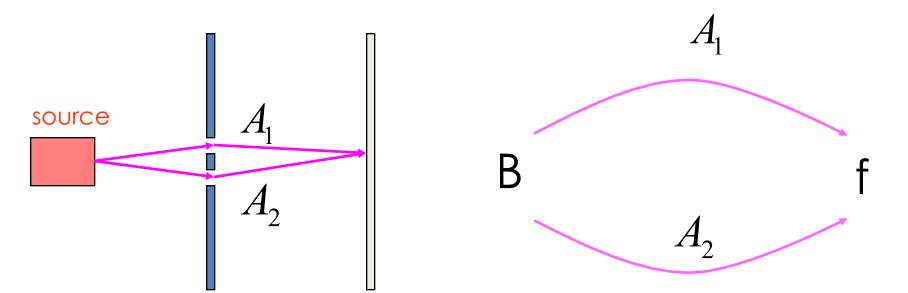
I∕/±

# families	# angles	# reducible phases	# irreducible phases
n	n(n-1)/2	2n-1	n(n+1)/2 –(2n-1)=(n-1)(n-2)/2
2	1		0
3	3		1
4	6		3

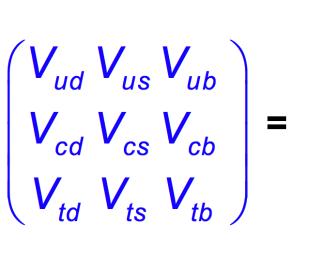
$$(u \quad c \quad t) \begin{pmatrix} V_{ud} & V_{us} & V_{ub} \\ V_{cd} & V_{cs} & V_{cb} \\ V_{td} & V_{ts} & V_{tb} \end{pmatrix} \begin{pmatrix} d \\ s \\ b \end{pmatrix} \qquad q_1 - \frac{1}{V_{q_1 q_2}} q_2$$

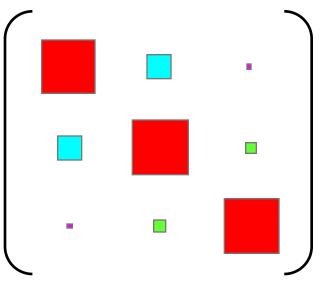
V_{CKM} Cabibbo-Kobayashi-Maskawa matrix

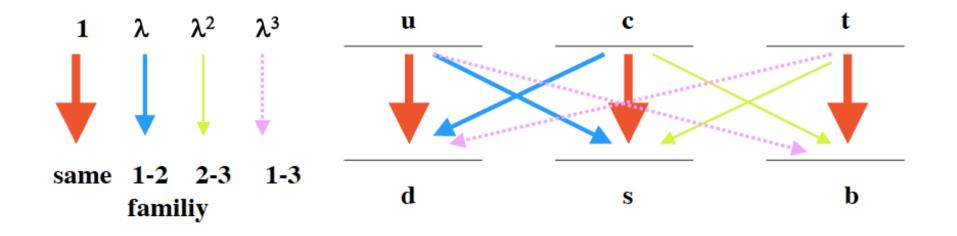




No prediction on the $V_{ij} \rightarrow$ they need to be measured \rightarrow Experimental observations :





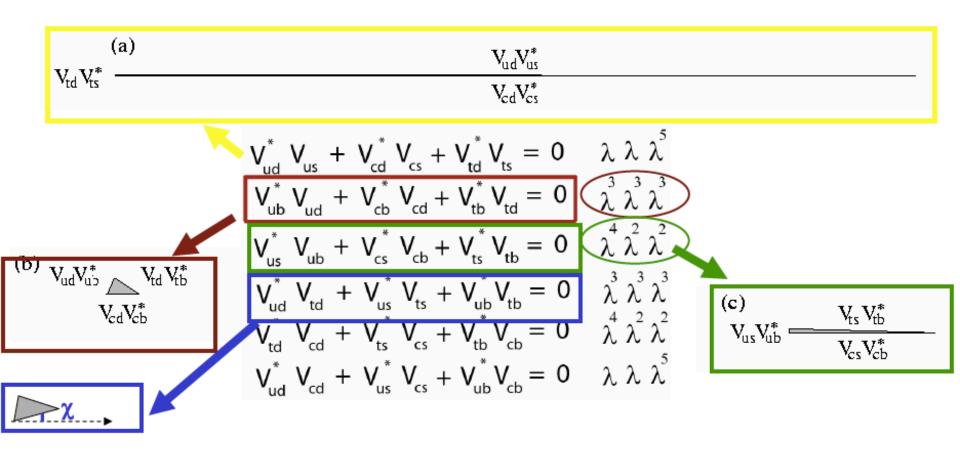


Measuring triangles

Stay within the 3 families

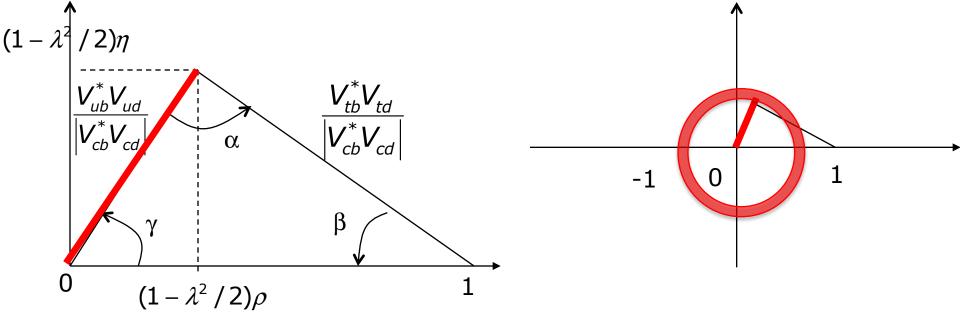
$$\begin{pmatrix} u & c & t \end{pmatrix} \begin{pmatrix} V_{ud} & V_{us} & V_{ub} \\ V_{cd} & V_{cs} & V_{cb} \\ V_{td} & V_{ts} & V_{tb} \end{pmatrix} \begin{pmatrix} d \\ s \\ b \end{pmatrix}$$
 Unitarity of $V_{CKM} \quad VV^{\dagger} = V^{\dagger}V = 1$
$$\Rightarrow 9 \text{ relations} \quad \sum_{k=1}^{n} V_{ik}V_{jk}^{*} = \delta_{ij},$$

The non-diagonal elements of the matrix products correspond to 6 triangle equations

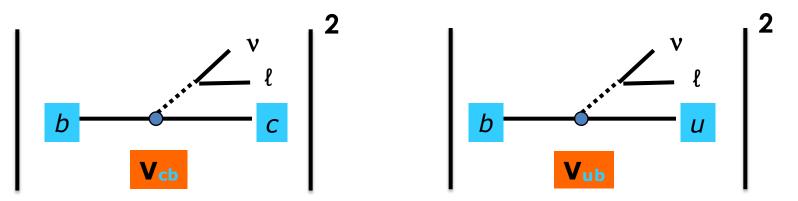


They all have the same area, proportionnal to the amount of CP violation in the SM

Measurements of the unitarity triangle parameters : some examples

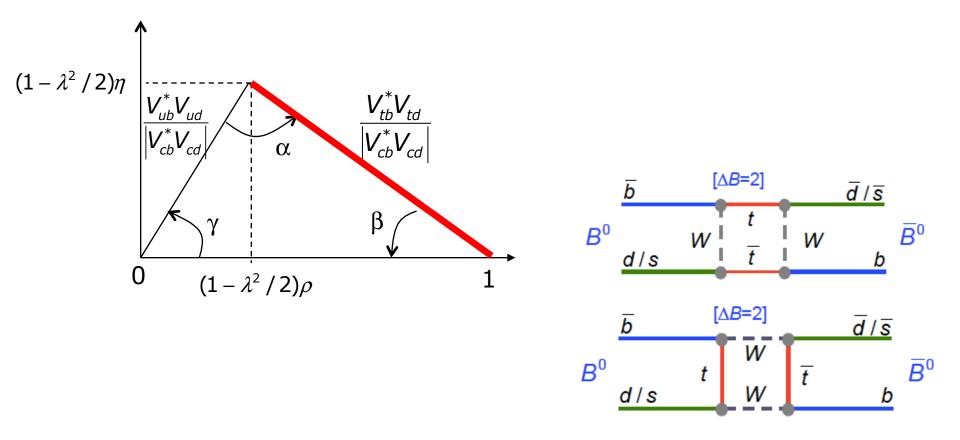


Rates of semileptonic B decays



Conceptually simple, complicated by QCD

The other side : $B^{0}-\overline{B}^{0}$ oscillations



Diagrams involving V_{td} or V_{ts}

The mixing phenomenon

Pairs of self-conjugate mesons that can be transformed to each other via flavour changing weak interaction transitions are:

$$|\mathbf{K}^{0}\rangle = |\overline{\mathbf{s}}\mathbf{d}\rangle$$
 $|\mathbf{D}^{0}\rangle = |\mathbf{c}\overline{\mathbf{u}}\rangle$ $|\mathbf{B}^{0}_{\mathbf{d}}\rangle = |\overline{\mathbf{b}}\mathbf{d}\rangle$ $|\mathbf{B}^{0}_{\mathbf{s}}\rangle =$

They are **flavour eigenstates** with definite quark content

useful to understand particle production and decay

Apart from the flavour eigenstates there are mass eigenstates:

- eigenstates of the Hamiltonian
- states of definite mass and lifetime
- They are propagating through space-time

$$\begin{vmatrix} B_{L} \rangle = p \begin{vmatrix} B^{0} \rangle + q \begin{vmatrix} \overline{B}^{0} \rangle \\ |B_{H} \rangle = p \begin{vmatrix} B^{0} \rangle - q \begin{vmatrix} \overline{B}^{0} \rangle \end{vmatrix}$$

Since flavour eigenstates are not mass eigenstates, the flavour eigenstates are mixed with one another as they propagate through space and time

 $\left| B^{0} \right\rangle$, $\left| \overline{B}^{0} \right\rangle$

bs

 $\ket{B_{\!\scriptscriptstyle L}}$, $\ket{B_{\!\scriptscriptstyle H}}$

$$|B_{H,L}(t)\rangle = e^{-i\left(M_{H,L}-i\frac{\Gamma_{H,L}}{2}\right)t}|B_{H,L}(t=0)\rangle + \frac{|B_L\rangle = p|B^0\rangle + q|\overline{B}^0\rangle}{|B_H\rangle = p|B^0\rangle - q|\overline{B}^0\rangle}$$

Time

The probability to observe a B^0 at time t if a B^0 was produced at time t=0 is :

$$\left|\left\langle B^{0}\left|H\right|B^{0}\left(t\right)\right
ight
angle ^{2}=rac{e^{-\Gamma t}}{2}\left(1+\cos\Delta mt
ight)$$

The probability to observe a B^0 at time t if a $\overline{B^0}$ was produced at time t=0 is :

$$\left|\left\langle \overline{B}^{0}\left|H\right|B^{0}\left(t\right)\right
angle \right|^{2}=rac{e^{-\Gamma t}}{2}\left(1-\cos\Delta mt
ight)$$

This is the mixing phenomenon!

evolution

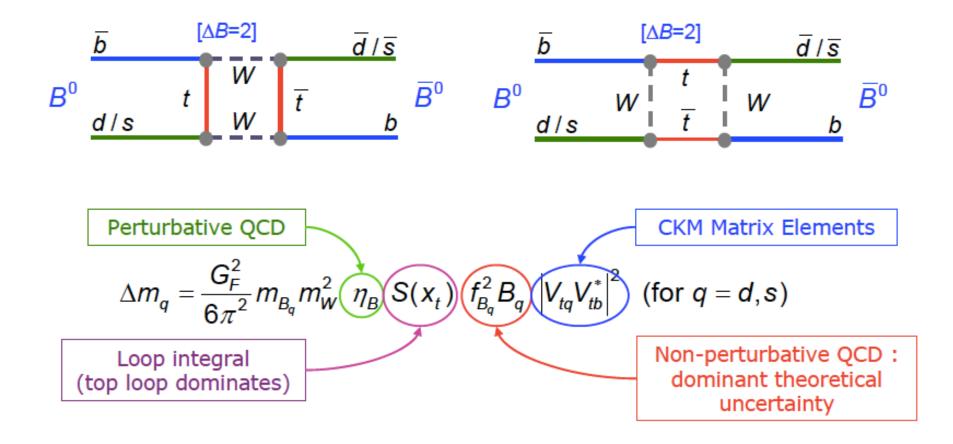
$$\frac{N_{Unmixed} - N_{Mixed}}{N_{Unmixed} + N_{Mixed}} \sim \cos \Delta m t$$

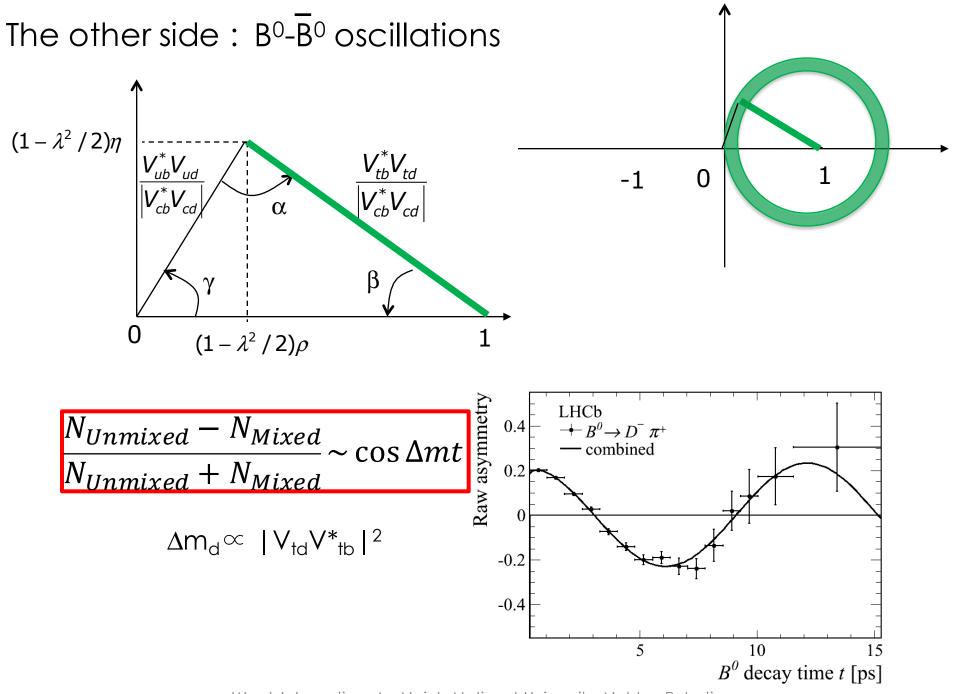
Simplified formulae assuming that the two mass eigenstates have the same lifetime and neglecting CP violation (q/p=1)

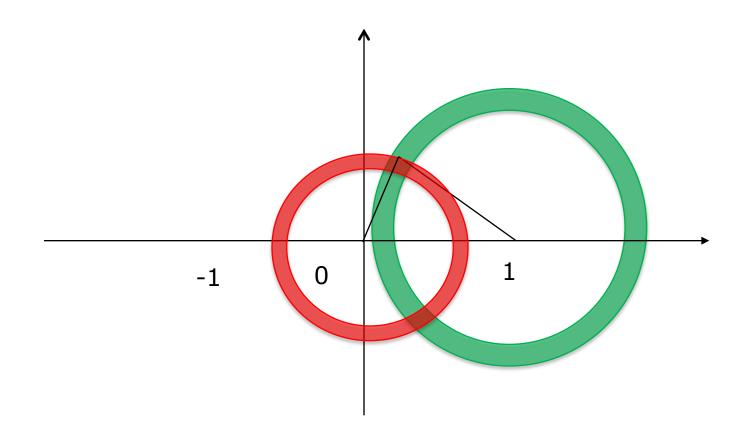
Let's come back to the unitarity triangle

Δm can be computed in the Standard Model

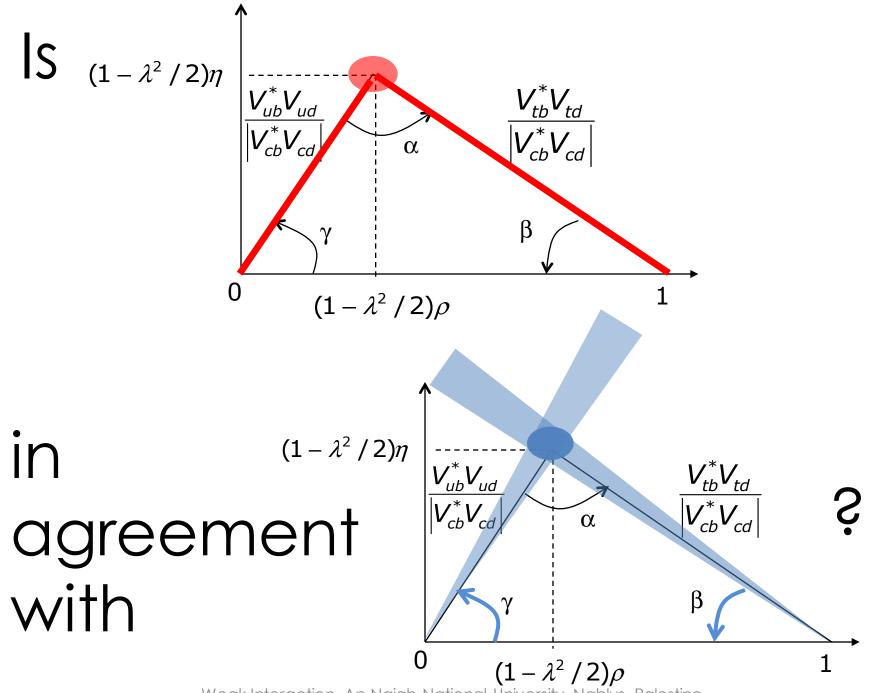
Effective FCNC Processes (CP conserving — top loop dominates in box diagram):







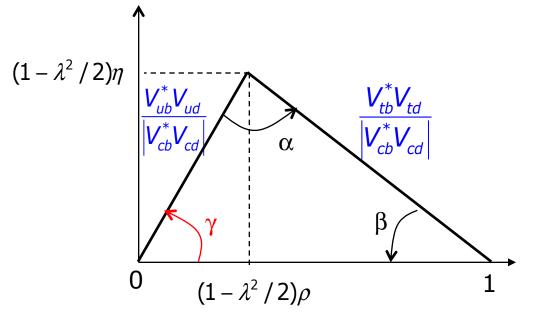
Are the two types of measurements compatible ?



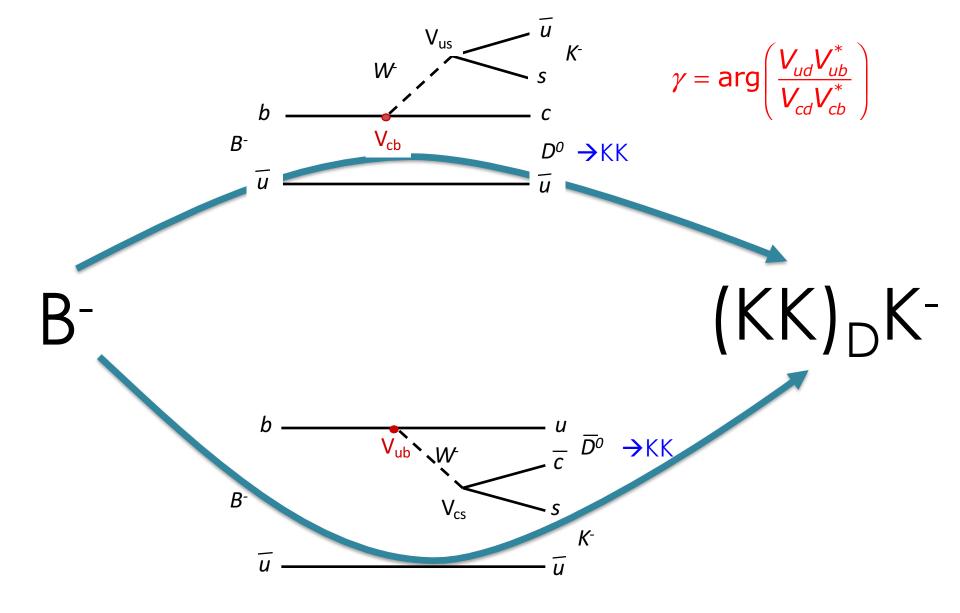
"the" unitarity triangle :

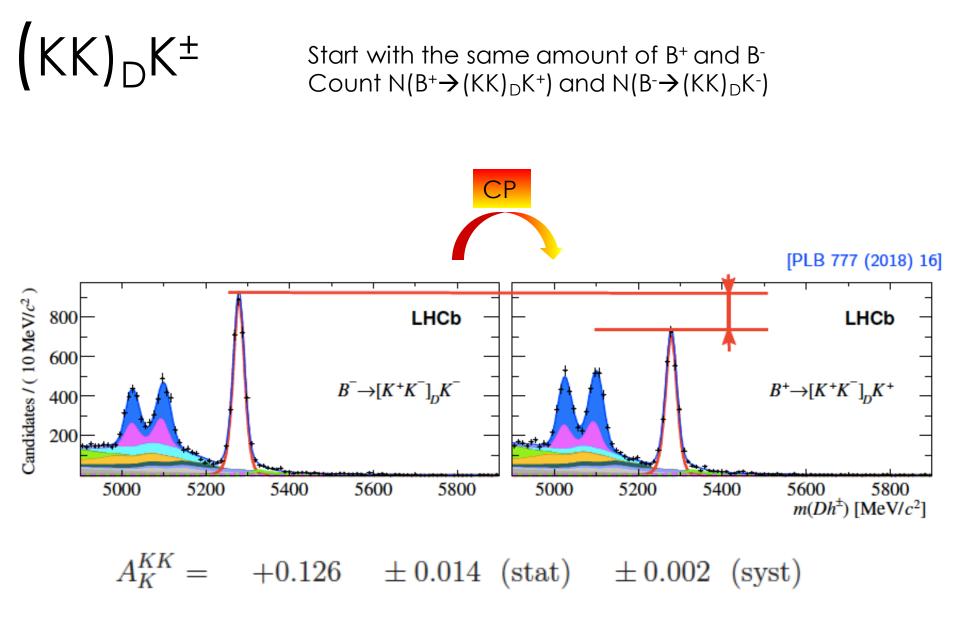
 $V_{ub}^{*}V_{ud} + V_{cb}^{*}V_{cd} + V_{tb}^{*}V_{td} = 0$

CP violation measurement

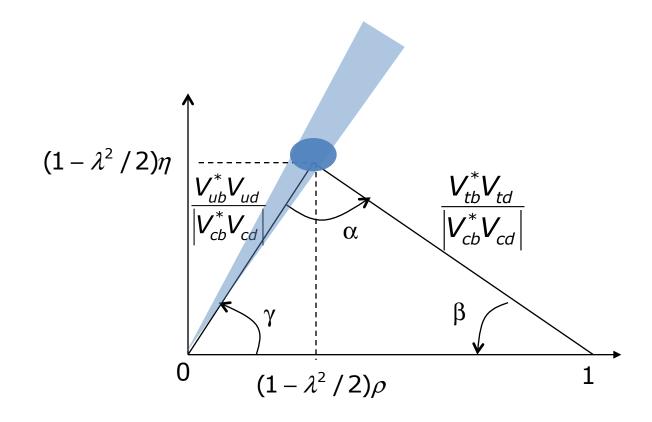


 $\gamma = \arg\left(\frac{V_{ud}V_{ub}^*}{V_{cd}V_{cb}^*}\right)$

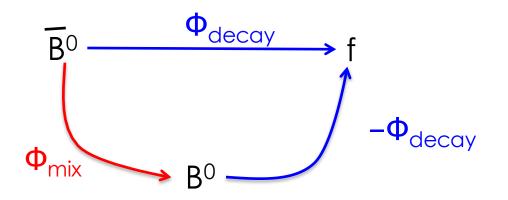




significantly different from 0!



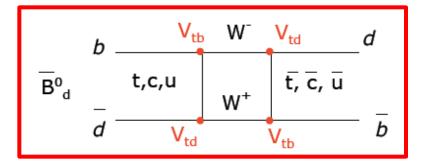
An example of CP induced by the interference between mixing and decay : the β angle

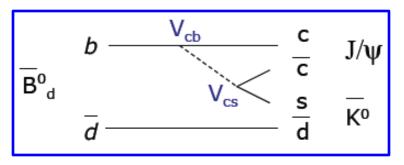


$$\Phi_{d} = \Phi_{mix} - 2 \Phi_{decay}$$





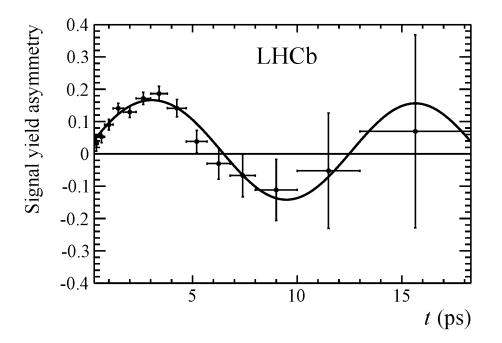


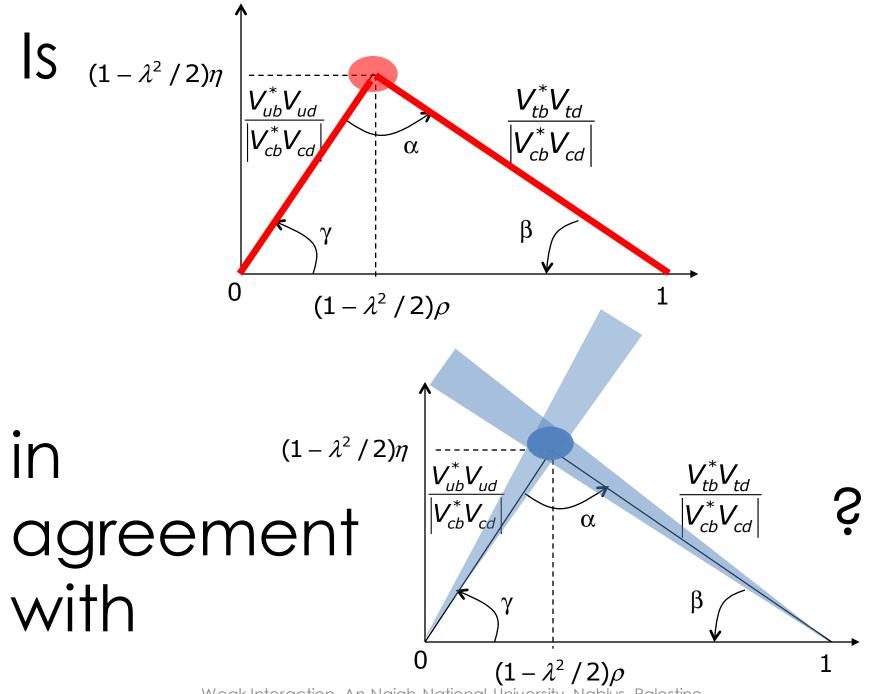


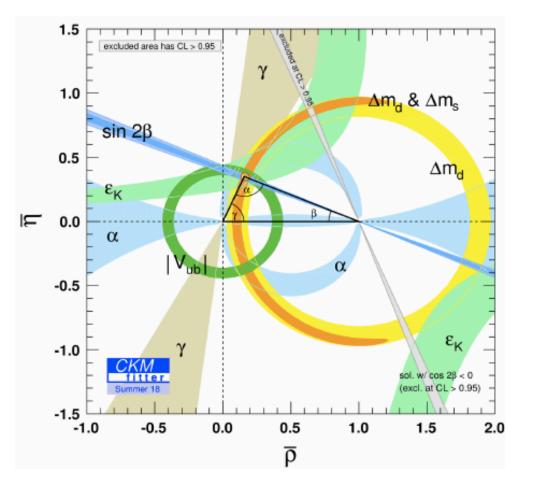
$$a_{f_{CP}}(t) = \frac{\operatorname{Prob}(B^{0}(t) \to f_{CP}) - \operatorname{Prob}(\overline{B^{0}(t)} \to f_{CP})}{\operatorname{Prob}(\overline{B^{0}(t)} \to f_{CP}) + \operatorname{Prob}(B^{0}(t) \to f_{CP})} =$$

 $= \sin(2\beta) \sin(\Delta mt)$

Pionnered by the B-factories





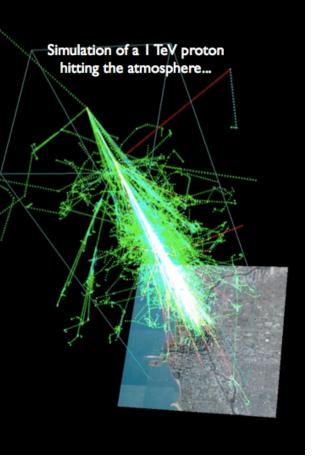


 $ar{
ho}=0.1577^{\,+0.0096}_{\,-0.0074}~(5\%~{
m unc.})\ ar{\eta}=0.3493^{\,+0.0095}_{\,-0.0071}~(2\%~{
m unc.})$

Sides and angles measurements in good agreement

The CKM model of CP violation has been confirmed

At the electroweak scale, the CKM mechanism dominates CP Violation

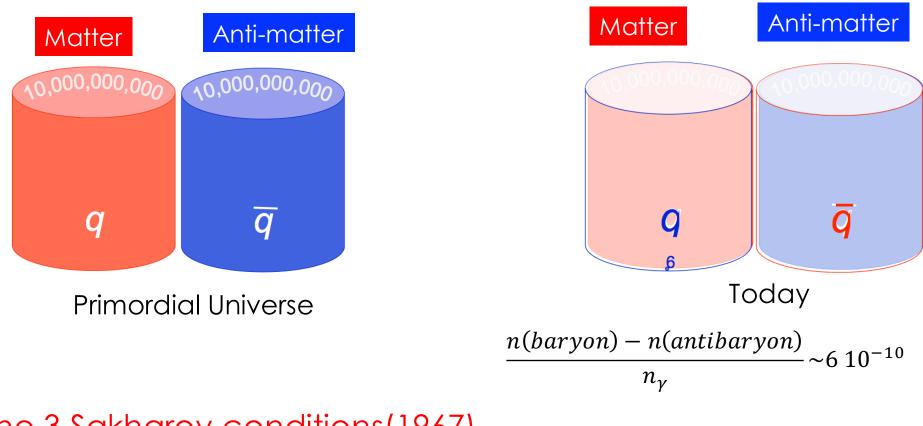


- Anti-matter in cosmic rays
- No sign of light emission (anti-galaxy ...)

•No significant sign of anti-nuclei (anti-He $^4 \ \dots$) Searches on-going



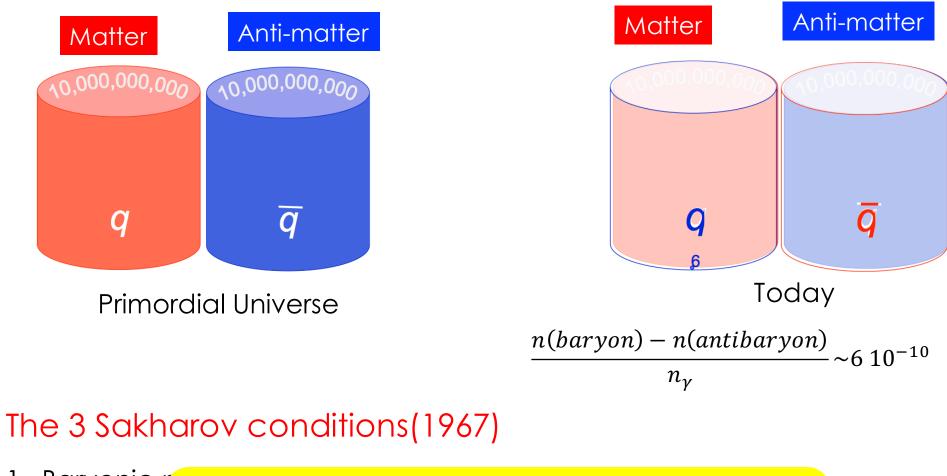
Anti-matter in the Universe and Big Bang



The 3 Sakharov conditions(1967)

- 1. Baryonic number violation: $X \rightarrow pe^{-1}$
- 2. C and CP symmetries violation: $\Gamma(X \rightarrow p e^{-}) \neq \Gamma(\overline{X} \rightarrow \overline{p} e^{+})$
- 3. To be out of equilibrium: $\Gamma(X \rightarrow p e^{-}) \neq \Gamma(p e^{-} \rightarrow X)$

Anti-matter in the Universe and Big Bang



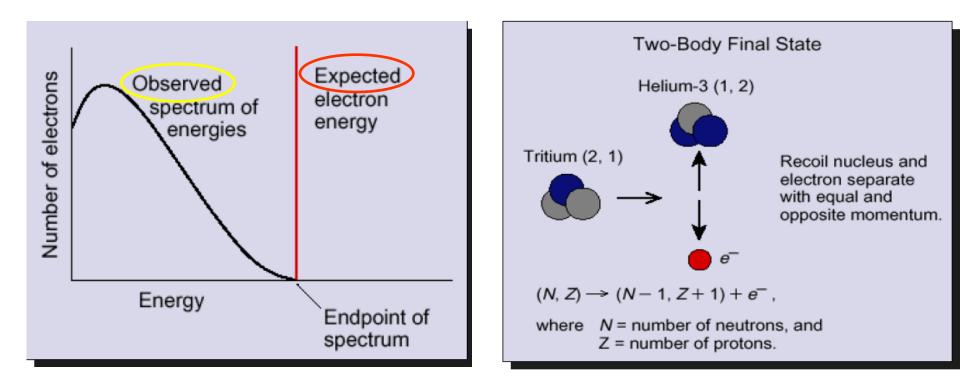
- 1. Baryonic r
- 2. C and CP

But the CP violation phase of the SM is orders of magnitude too small

3. To be out or equilibrium. It a pe price of

Neutrinos

The β decay

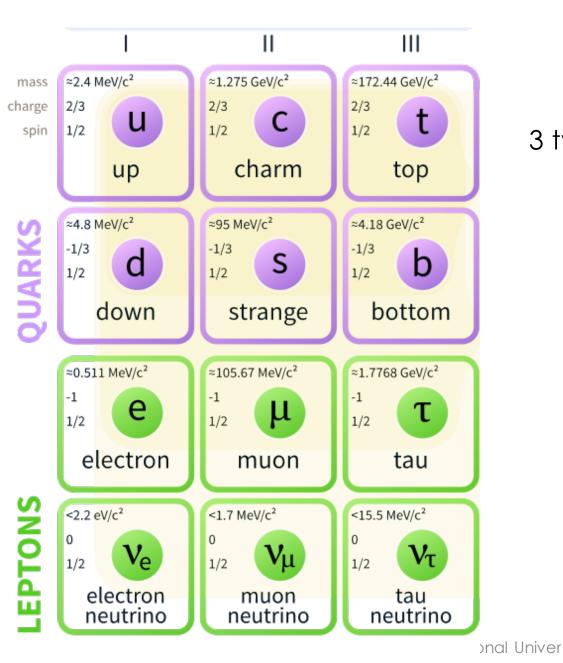


If β decay proceeds through $n \rightarrow p e^{-}$ the energy conservation predicts a monochromatic spectrum

1914 Chadwick observes a continuous spectrum ...

→ Energy conservation is violated (Bohr) or an other particle is in the game (Pauli) : β decay : n→p e- \overline{v}_e

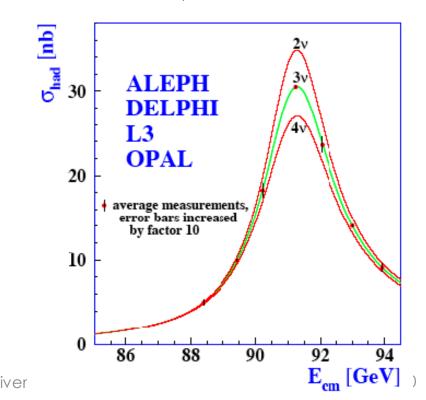
Fermionic sector of the SM :

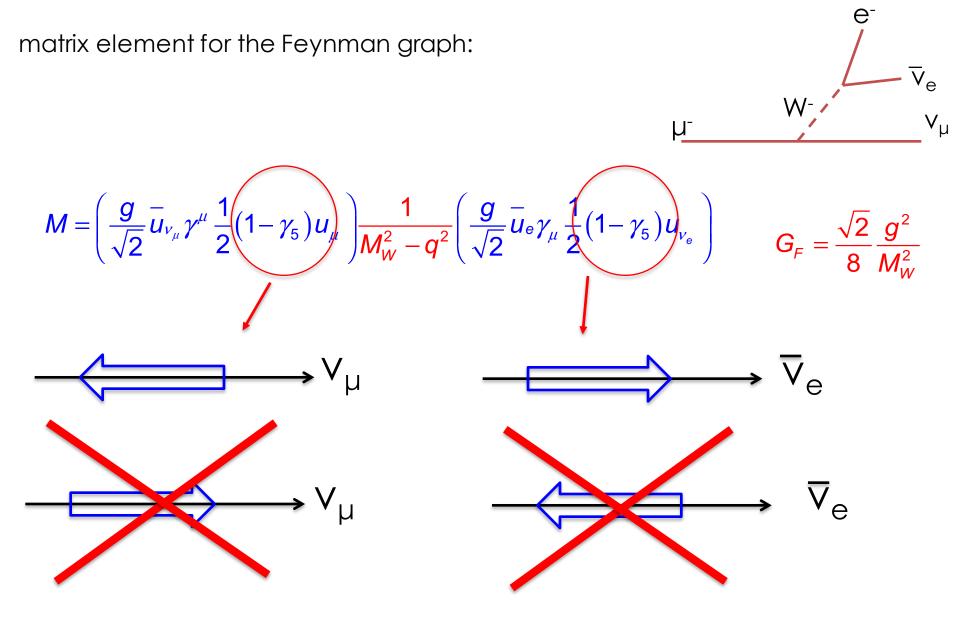


Lepton number conservation

3 types of neutrinos with m = 0

 $N_v = 2.9840 \pm 0.0082$

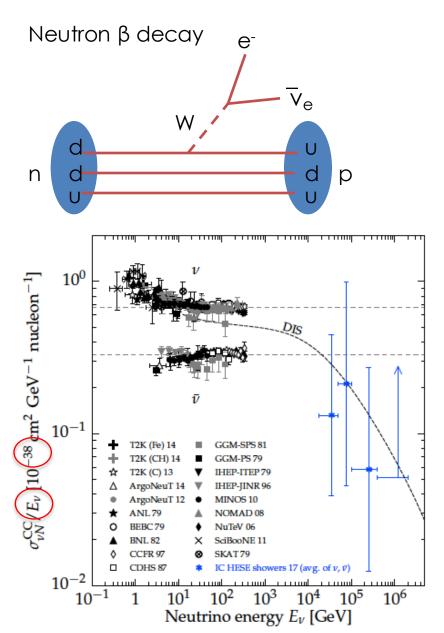


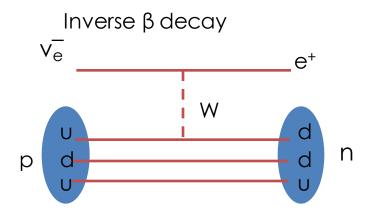


The v is left handed (the anti-neutrino is right handed)

How to detect a neutrino

If v are produced by β decay, they can be detected using the inverse reaction.





 $\sigma(\nu\,p)\sim 10^{\text{-}43}\,cm^2$ for E_ν = 3 MeV

So one needs :

- Intense neutrino sources
- Large mass detectors

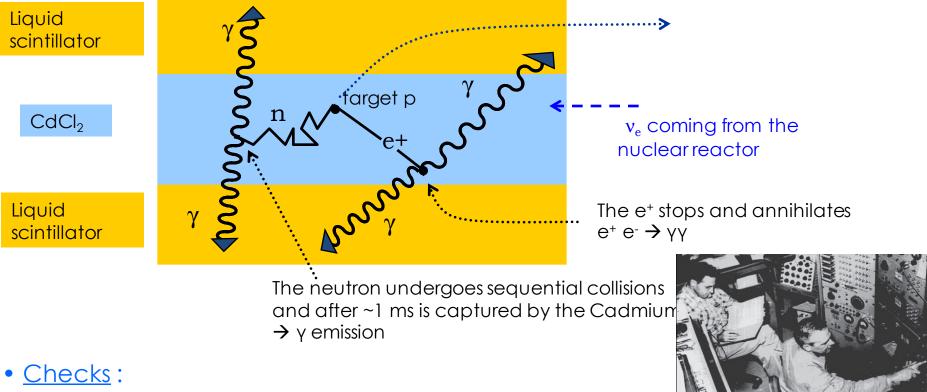
Sun

Cosmic rays interactions Reactors Accelerators

Jational University, Nablus, Palestine

Direct experimental evidence of ve

- Reminder: around 1930, Pauli and Fermi made the hypothesis of the ν_e
- In 1956, Reines-Cowan experiment : experimental evidence using a nuclear reactor : search for the $\bar{v}_e p \rightarrow n e^+$ reaction $\bar{v}_e p \rightarrow n e^+$

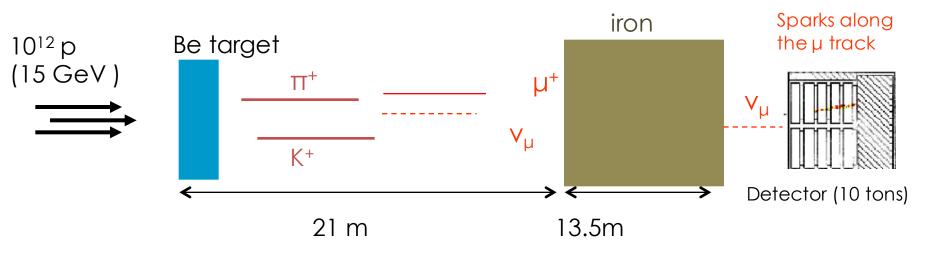


- Target without cadmium
- data taking with nuclear reactor OFF

 \rightarrow 3.0 ±0.2 signal events/hour

Two other types de neutrinos

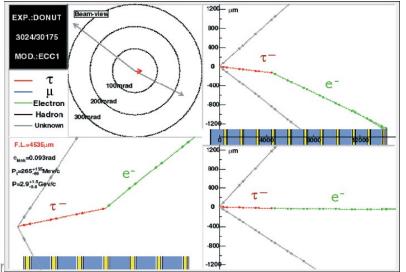
• In 1962 : Schwartz, Lederman et Steinberger experiment at BNL : $V_{\mu} \neq V_{e}$



• in 2000 the third neutrino v_{τ} (DONUT) at Fermilab :

$$D_s \rightarrow \tau v_{\tau}$$
 $v_{\tau} + N \rightarrow \tau + X$
 v_{τ} beam

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3 years of data taking : 4 unambiguous v_{\tau} signal events
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BUT:

In a deep mine the Davis Chlorine experiment

Inverse beta decay Cl³⁷+v - Ar³⁷+e

Searching for v_e

Ar is chemically very different from Chlorine \rightarrow can be separated

It is radioactive and reverts to Cl³⁷ emitting an Auger electron (lifetime 35 days)

Count a few atoms in a tank of few hundred tons !



The observed rate was about 3 times smaller than the predicted one ...

Which one is wrong ? The experiment or the Solar Model ?

Neutrinos mixing

If neutrinos have masses : Mass eigenstates ≠ weak (flavour) eigenstates

→ 3x3 matrix (CKM-style)

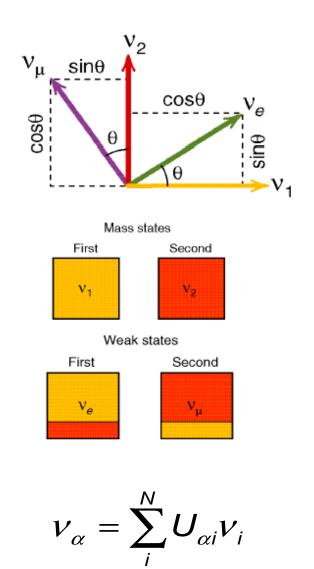
Assume for simplicity two families : the 2 basis are connected through a simple rotation :

$$\begin{pmatrix} v_e \\ v_\mu \end{pmatrix} = \begin{pmatrix} \cos\theta & \sin\theta \\ -\sin\theta & \cos\theta \end{pmatrix} \begin{pmatrix} v_1 \\ v_2 \end{pmatrix} = U \begin{pmatrix} v_1 \\ v_2 \end{pmatrix}$$

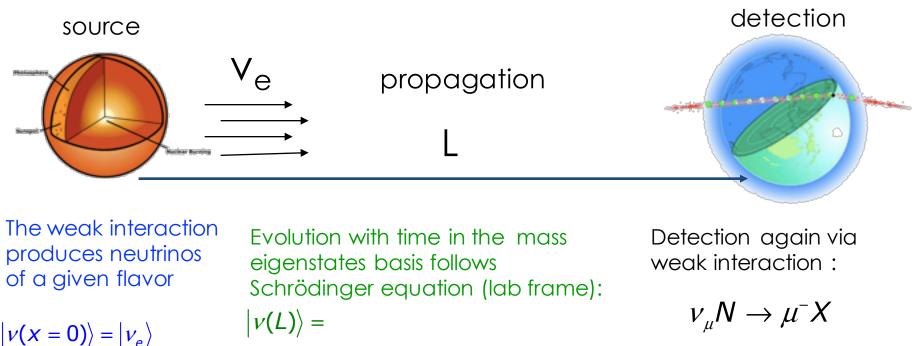
In the case of 3 families :

$$\begin{pmatrix} \boldsymbol{v}_{e} \\ \boldsymbol{v}_{\mu} \\ \boldsymbol{v}_{\tau} \end{pmatrix} = \begin{pmatrix} \boldsymbol{U}_{e1} \boldsymbol{U}_{e2} \boldsymbol{U}_{e3} \\ \boldsymbol{U}_{\mu 1} \boldsymbol{U}_{\mu 2} \boldsymbol{U}_{\mu 3} \\ \boldsymbol{U}_{\tau 1} \boldsymbol{U}_{\tau 2} \boldsymbol{U}_{\tau 3} \end{pmatrix} \begin{pmatrix} \boldsymbol{v}_{1} \\ \boldsymbol{v}_{2} \\ \boldsymbol{v}_{3} \end{pmatrix}$$

Weak interaction eigenstates (n_a) PMNS matrix (Pontecorvo, Maki, Nakagawa, Sakata) Mass eigenstates (n_i)



Neutrinos oscillations : the case for 2 families



 $= \cos\theta |v_1\rangle + \sin\theta |v_2\rangle$

 $|v(L)\rangle = e^{-i(E_1t - p_1L)} \cos \theta |v_1\rangle + e^{-i(E_2t - p_2L)} \sin \theta |v_2\rangle$

Ultra relativistic neutrinos of momentum p

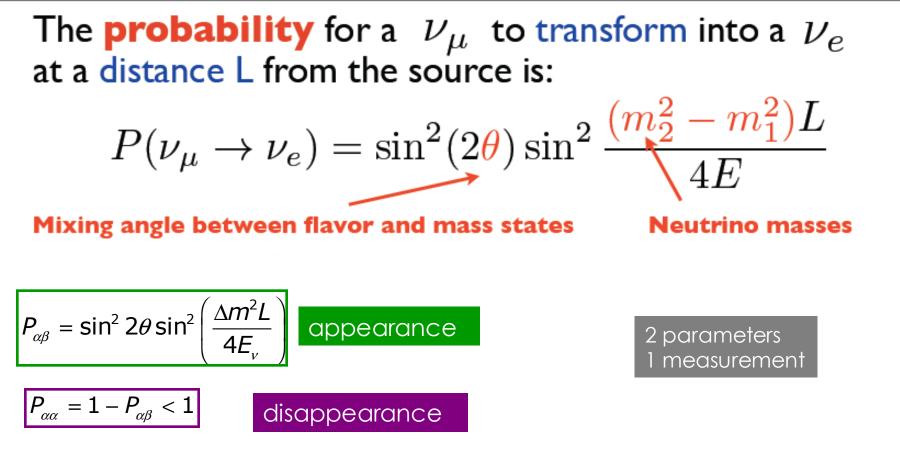
$$p \gg M \implies E_i = \sqrt{p^2 + M_i^2} \approx p + \frac{M_i^2}{2p}$$

$$\left|v(L)\right\rangle = e^{-i\frac{M_{1}^{2}}{2p}L}\cos\theta\left|v_{1}\right\rangle + e^{-i\frac{M_{2}^{2}}{2p}L}\sin\theta\left|v_{2}\right\rangle$$

 $P(v_e \rightarrow v_{\mu}) = |\langle v_{\mu} | v(L) \rangle|^2$ $\approx \sin^2 2\theta \sin^2 \frac{\Delta m_{12}^2 L}{4E}$ $\Delta m_{12}^2 = M_1^2 - M_2^2$

 $v_e N \rightarrow e^- X$

E is the average energy of the mass eigenstates



The other flavour eigenstates (not present at t=0) appear during the propagation

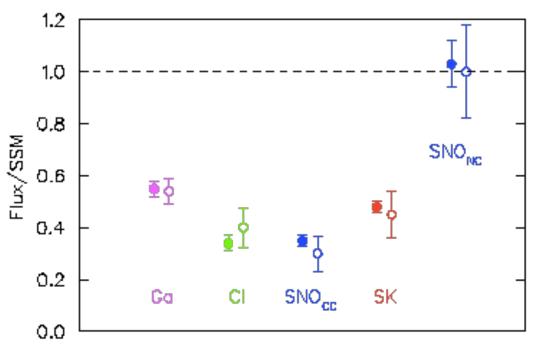
The total flux is conserved.

Neutrinos oscillations are only sensitive to the difference of the masses squared (NOT to the absolute value)

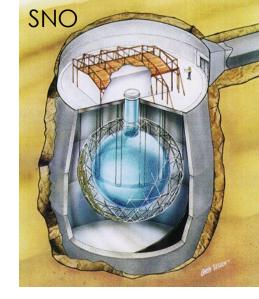
The mixing phenomenom can be tested using :

- Neutrinos from the sun
- Neutrinos from reactors
- Neutrinos from accelerators

If neutrinos oscillation is observed it implies that neutrinos have masses ...



Neutrinos from the sun !



the neutral current reaction (NC) measures the ($v_e + v_\mu + v_{\tau,}$) flux

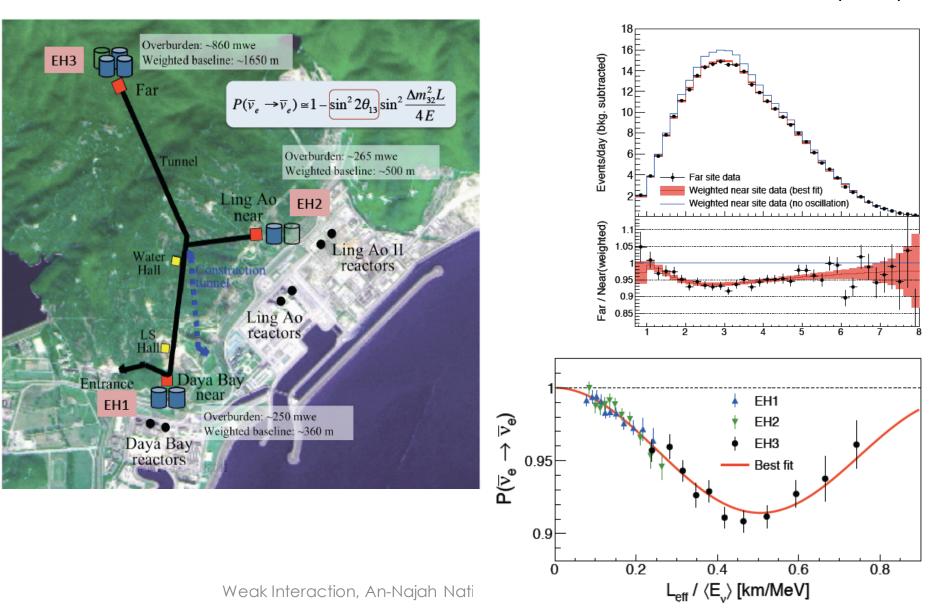
The charged current (CC) one the ν_{e} flux

The combination of all solar neutrino experiments (before SNO) implied that solar neutrinos were disappearing between production (in the sun core) and detection in the earth.

So the sun is shining the expected number of neutrinos but many of them are detected as and v_{μ} and/or $v_{\tau,}!$

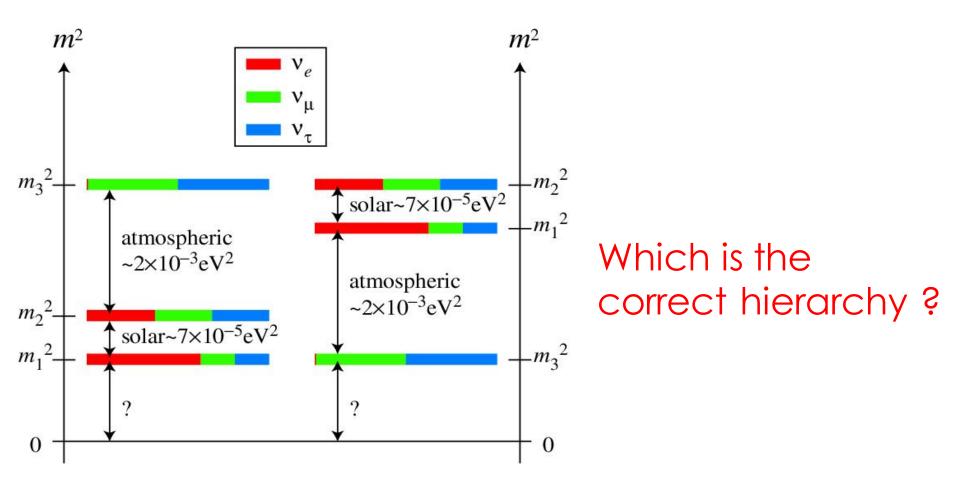
Neutrinos from the reactors : \overline{v}_{e}

PRL 115 111802 DayaBay

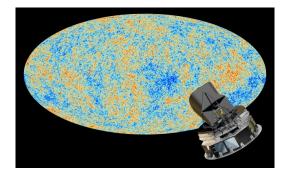


What is measured : $m_2^2 - m_1^2$ and $m_3^2 - m_2^2$

→ the neutrinos have masses but two scenarii are possible :



Cosmological constraints on $\Sigma m(v)$



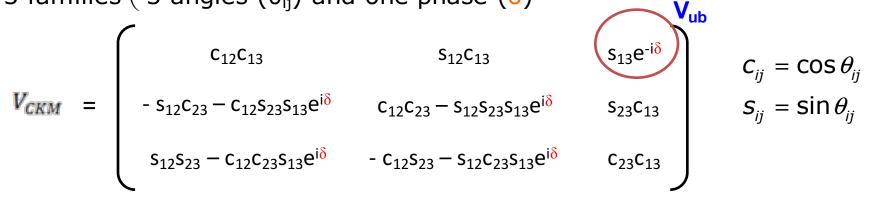
- The neutrinos mass would modify the (delicate) balance between gravity and the Hubble expansion
- They would also have also effect on the structure formation ...
- Small modifications in the Cosmological Microwave Background which is the fingerprint of what happended at the very beginning of the Universe

 $\rightarrow \Sigma m(v) < \sim 0.2 eV$

Weak interaction in summary

- All quarks and leptons are sensitive to the weak interaction
- M_W~M_Z~100 GeV
 - \rightarrow short range
 - → Extremely weak : (~ 10⁻⁸ smaller intensity than the strong interaction at a distance of 1 fm)
- The weak interaction
 - violates maximally C and P
 - does not conserve the flavour
 - Exhibits a tiny CP violation
- The weak and mass eigenstates of quarks are not the same, they are related via V_{CKM} which is a natural source of CP violation
- Neutrinos are only sensitive to the weak interaction
 - They have masses ! (but we do not know the values)
 - They may be an open window to physics beyond the Standard Model

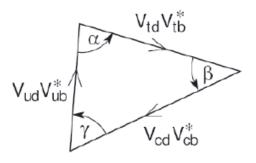
3 families (3 angles (θ_{ij}) and one phase (δ)



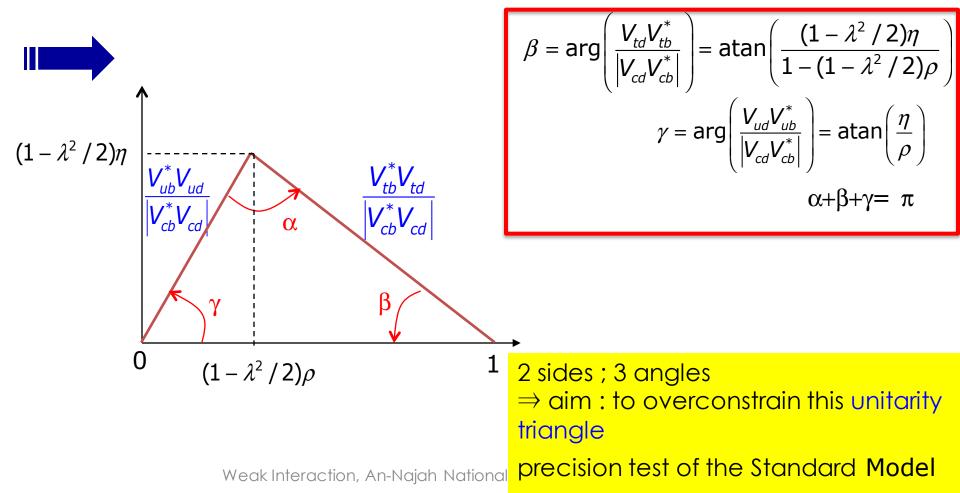
→ Parametrization in power of λ (=sin θ_c) = s₁₂ = |V_{us}| ~ 0.22

$$\begin{pmatrix} 1 - \lambda^{2} / 2 & \lambda & A\lambda^{3} (\rho - i\eta) \\ -\lambda & 1 - \lambda^{2} / 2 & A\lambda^{2} \\ A\lambda^{3} (1 - \rho - i\eta) & -A\lambda^{2} & 1 \end{pmatrix} + \vartheta(\lambda^{4}) \qquad \begin{array}{l} \lambda = \sin \theta_{c} \sim 0.22 \\ A \sim 0.80 \\ \rho \sim 0.20 \\ \eta \sim 0.35 \end{array}$$

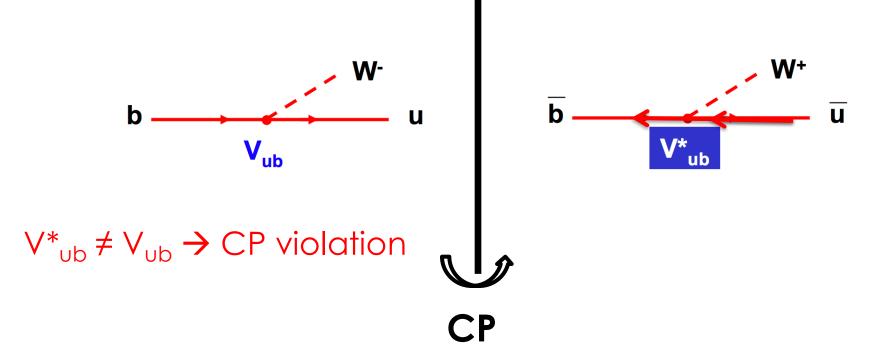
"the" unitarity triangle :
$$V_{ub}^*V_{ud} + V_{cb}^*V_{cd} + V_{tb}^*V_{td} = V_{td}V_{tb}^* = A\lambda^3(1 - \rho - i\eta) + A\lambda^5(\rho + i\eta)$$
$$V_{ud}V_{ub}^* = A\lambda^3(\rho + i\eta) \times (1 - \frac{\lambda^2}{2}) \qquad \text{at order } \lambda^5$$
$$V_{cd}V_{cb}^* = -A\lambda^3$$



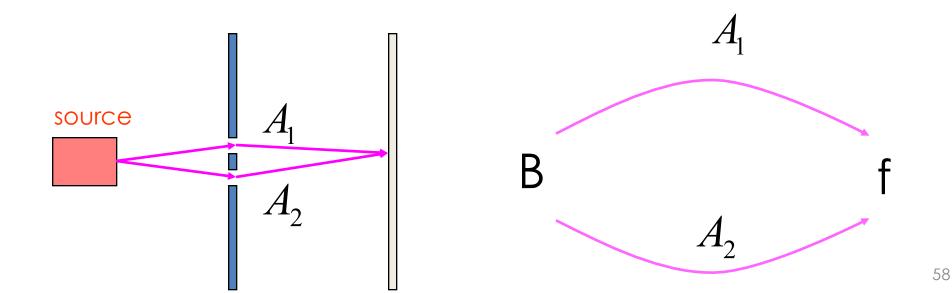
Basis of the triangle aligned on the real axis, normalized to 1



CP violation



If you just have one amplitude : no sensitivity on phase $(|V_{ij}|^2 = |V_{ij}^*|^2)$



Let's come back to the unitarity triangle

